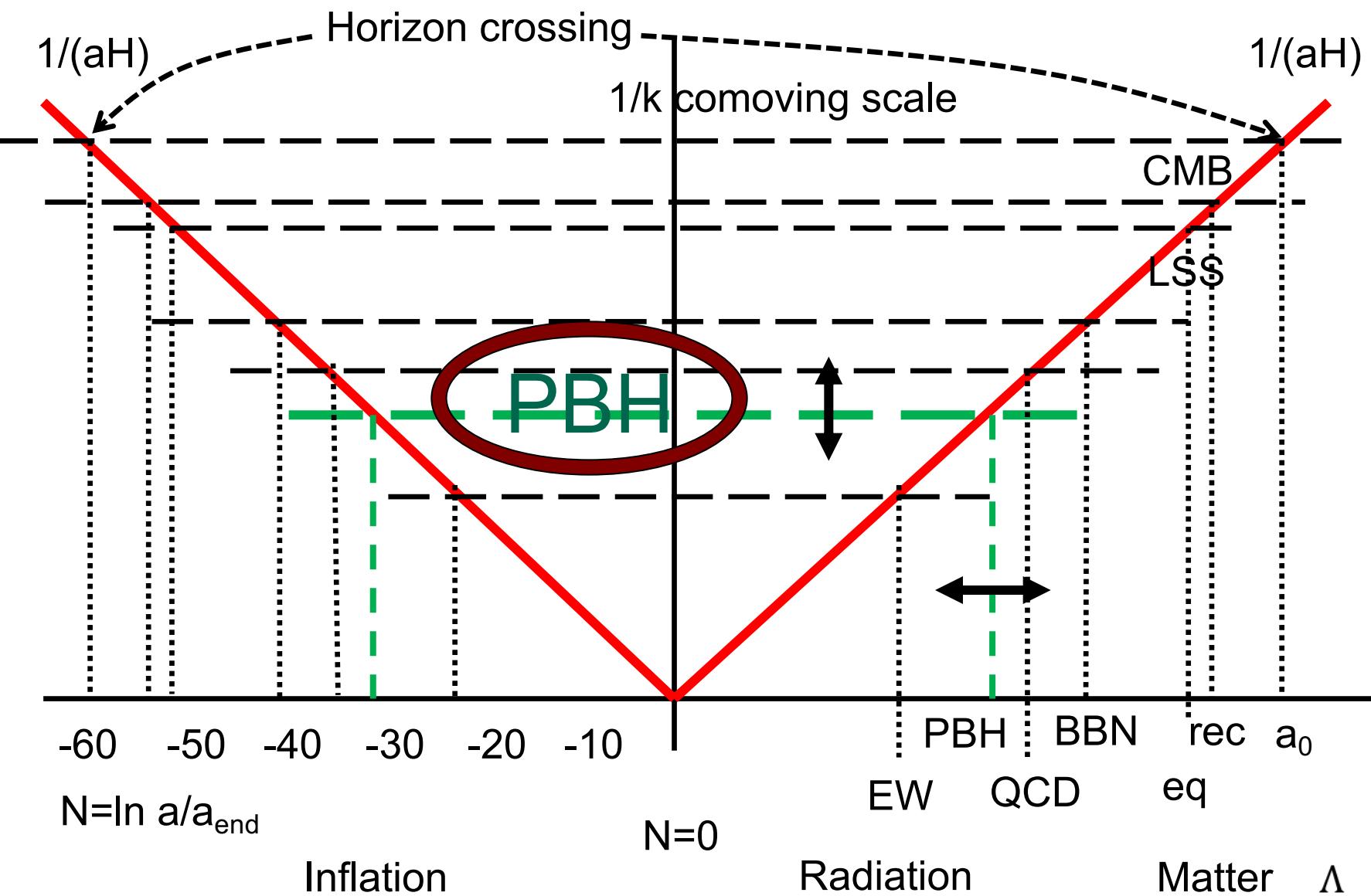


Have we detected Primordial Black Holes at LIGO-Virgo?

LVEM Zoom Meeting, 22nd October 2019

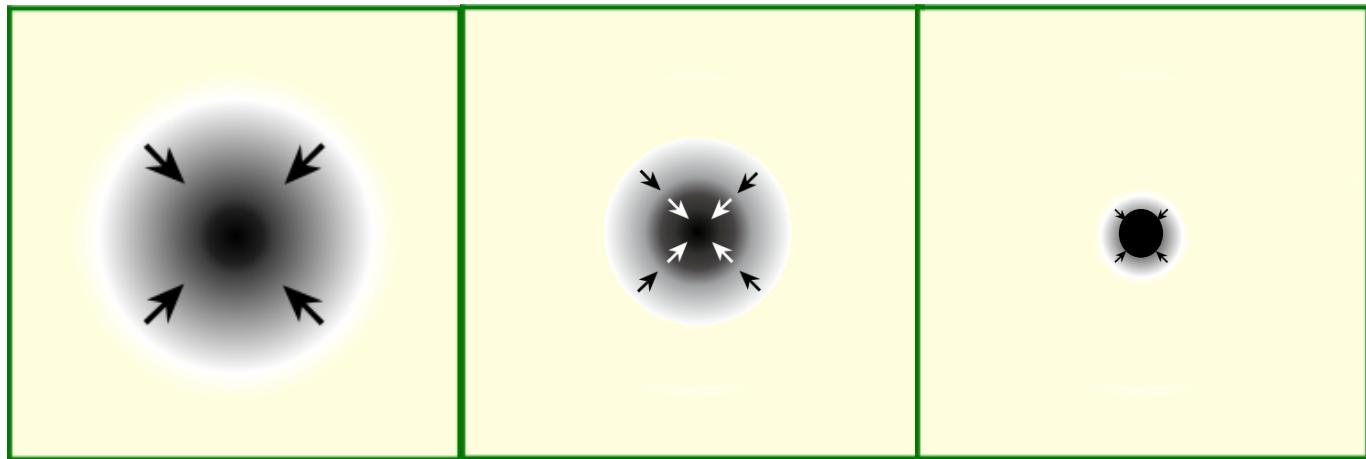
Juan García-Bellido
IFT-UAM/CSIC Madrid

Inflation

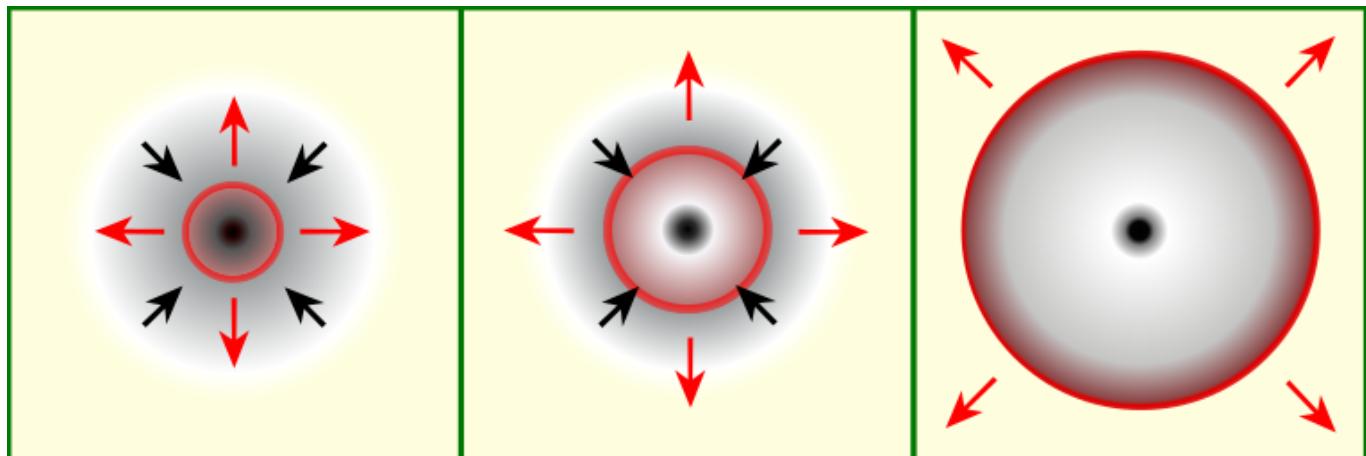


Gravitational Collapse

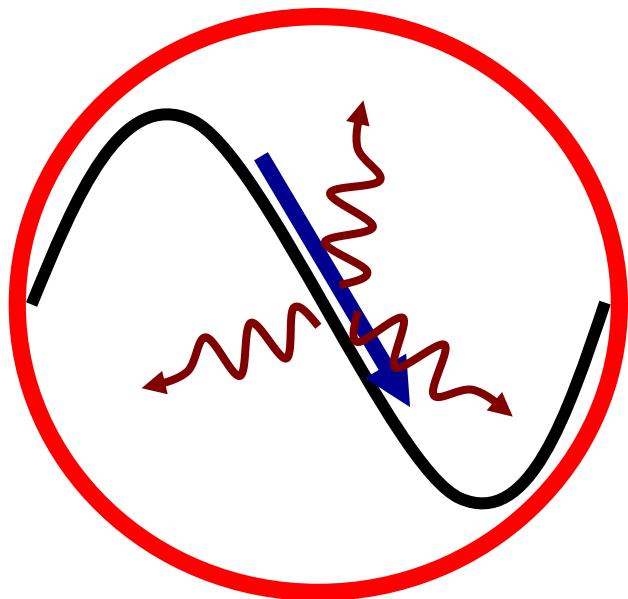
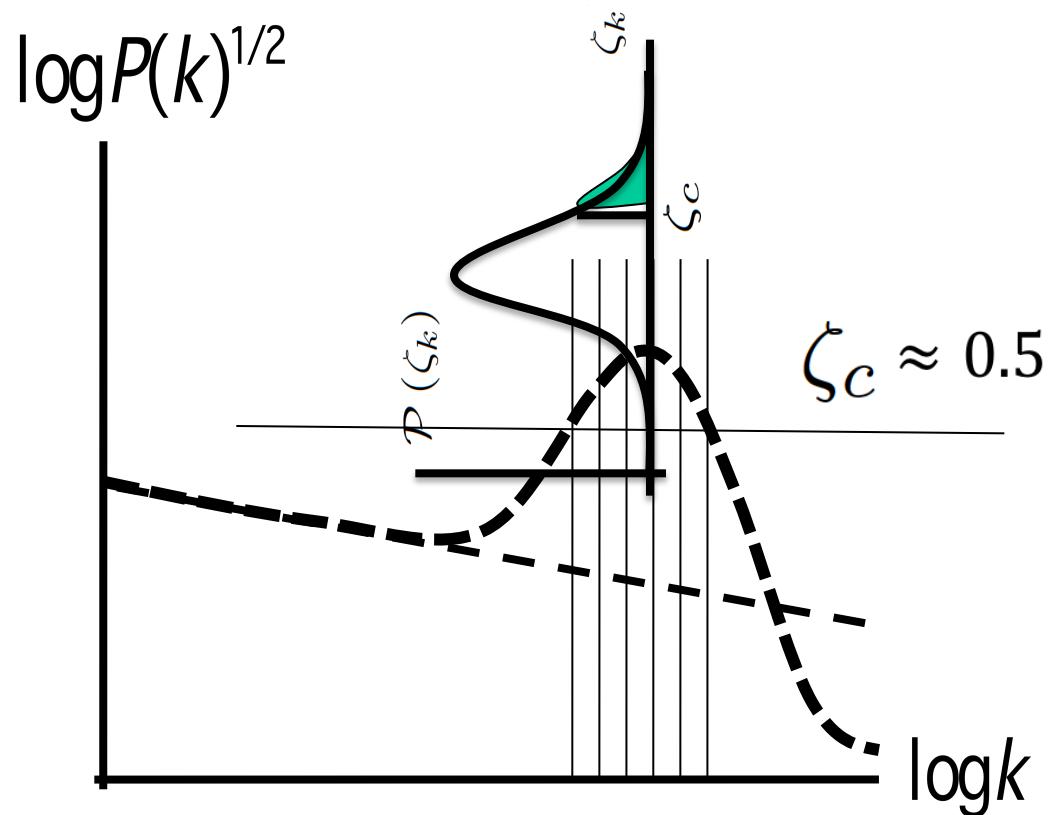
Gravity wins



Radiation wins



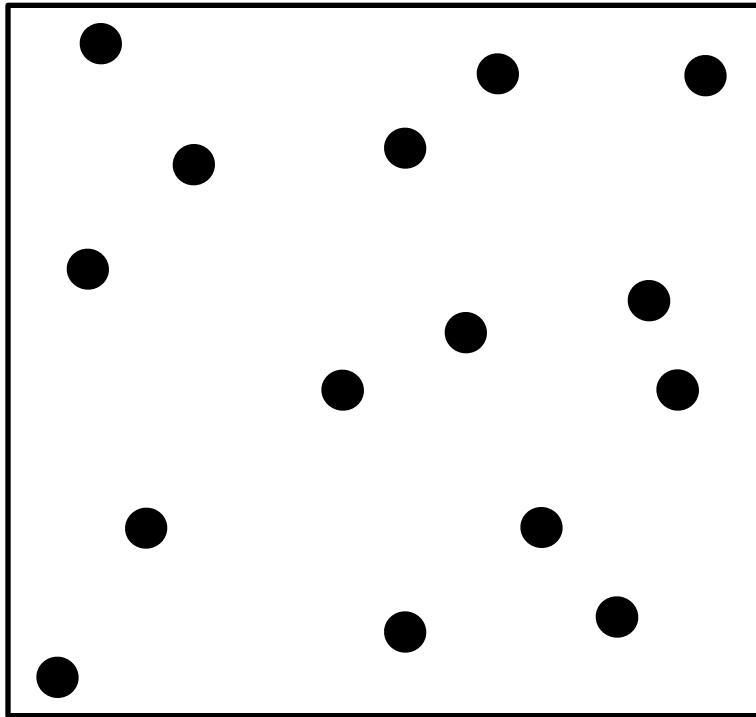
Gravitational Collapse of PBH



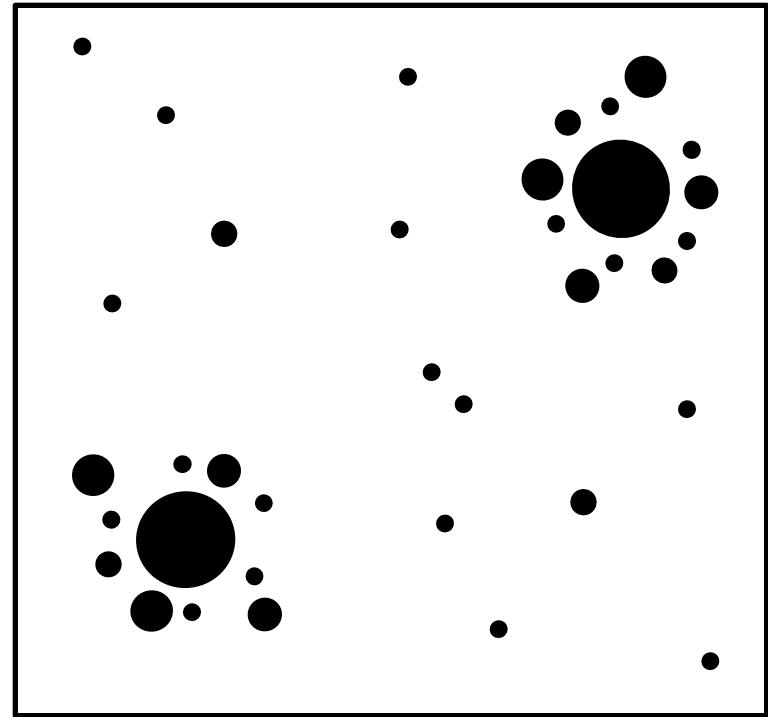
$$M_{PBH} \approx 30 M_\odot e^{2(N-36)}$$

Inflationary predictions

- Wide mass distribution
- Clusters of PBH: $N_{\text{cl}} \sim 100-1000$, comoving size $\sim 1 \text{ mpc}$



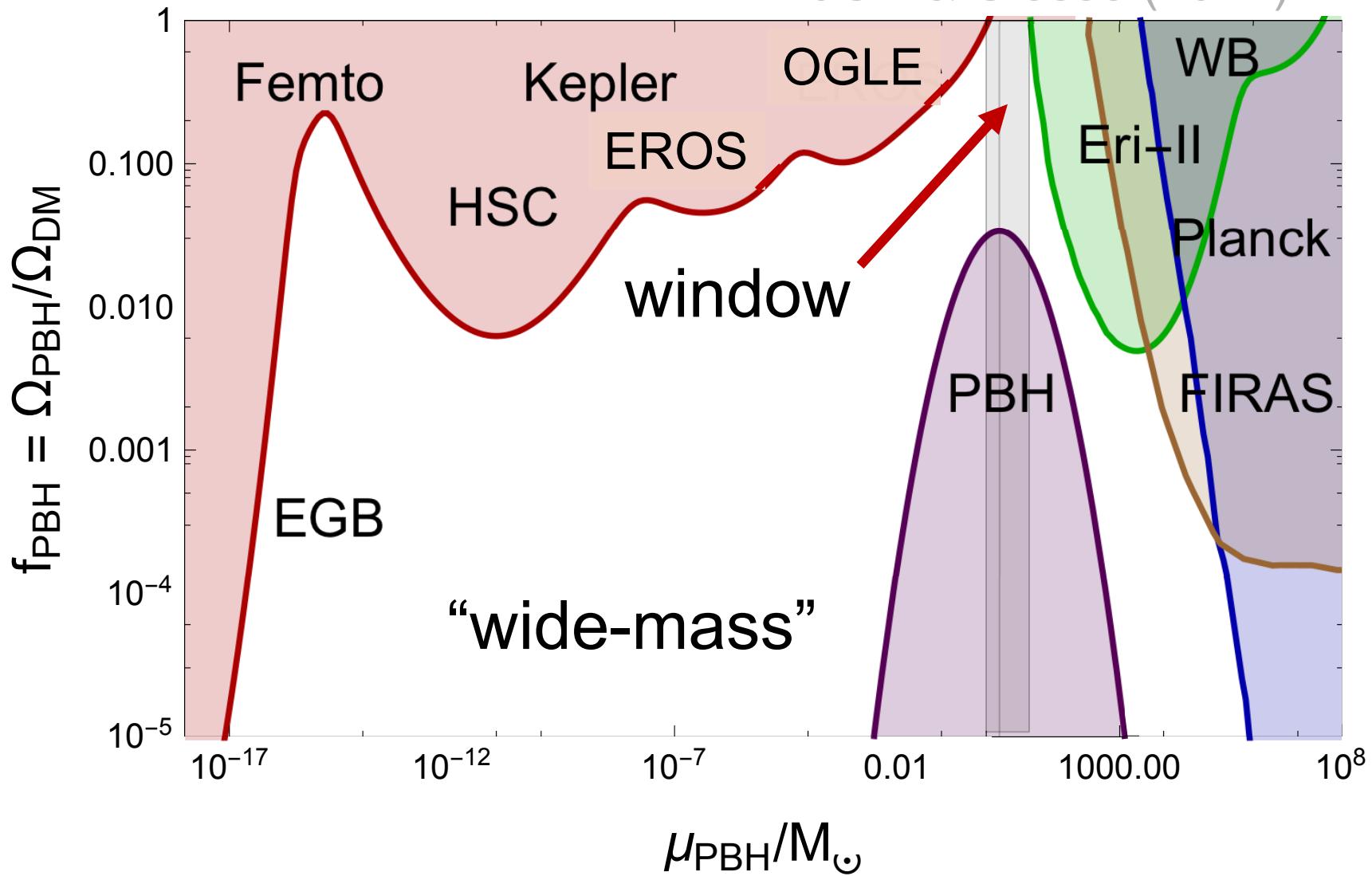
uniform single-mass
is already ruled out



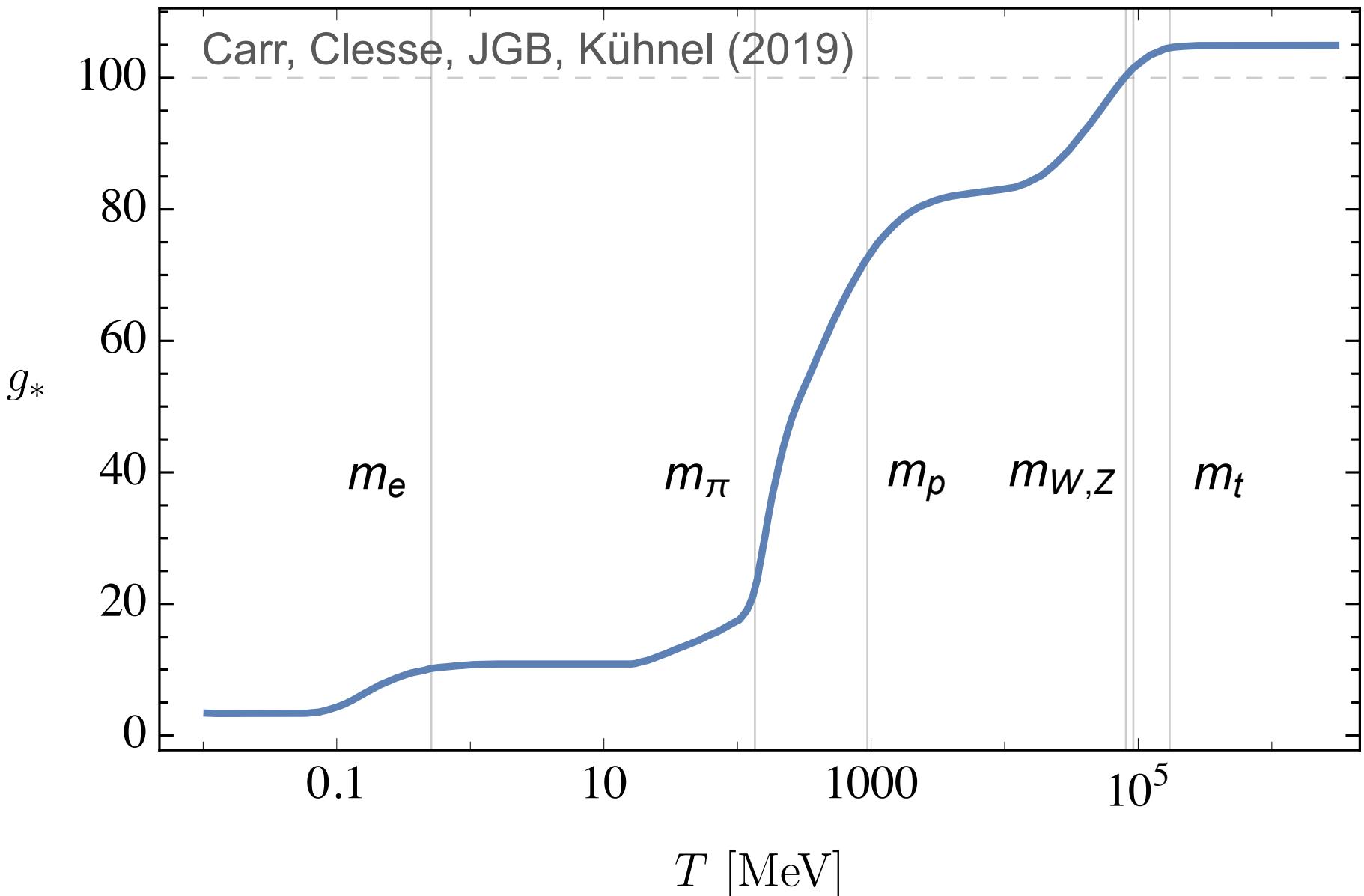
clustered wide-mass
is still viable

PBH constraints

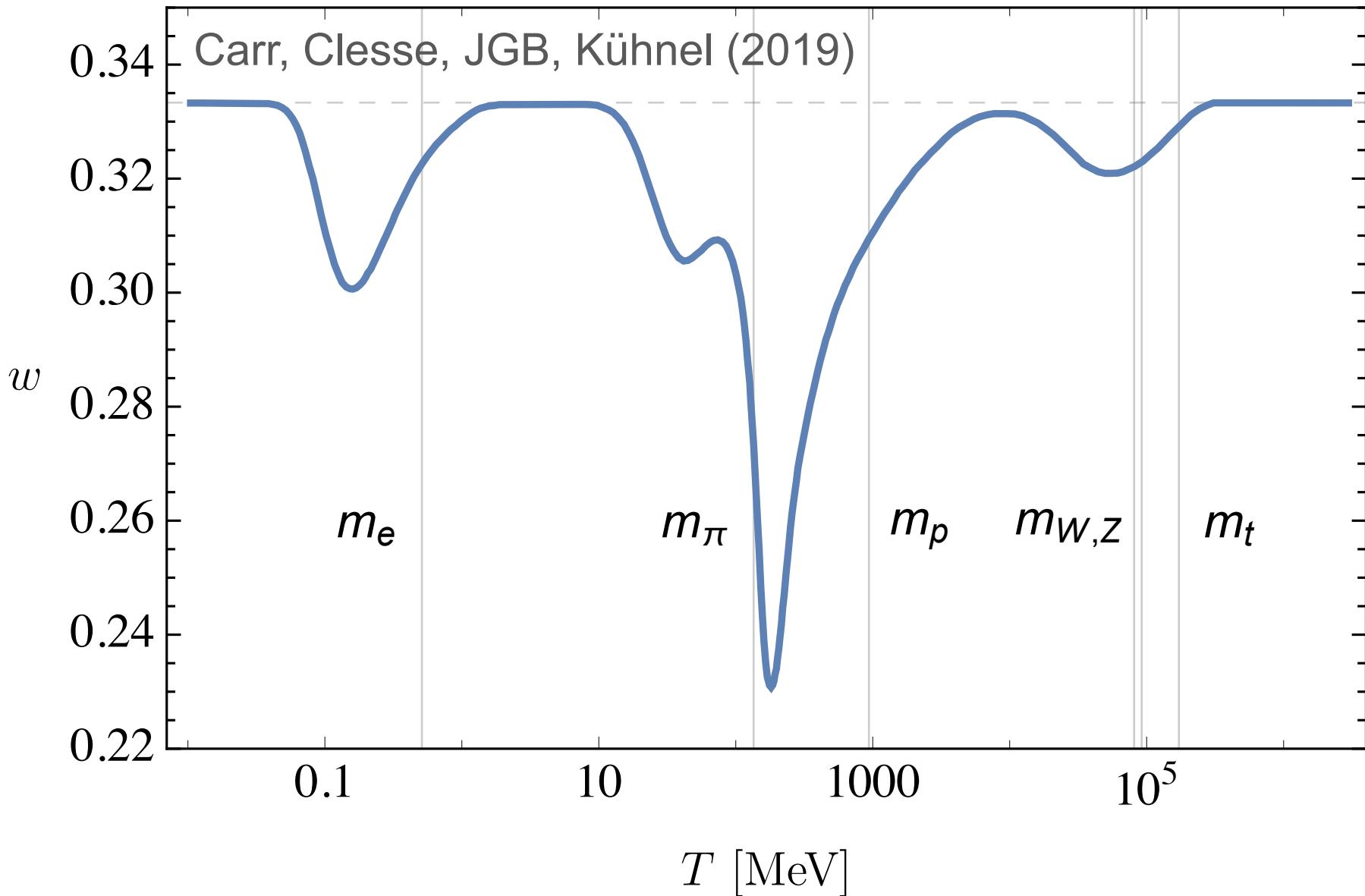
JGB & Clesse (2017)



Thermal history of the universe

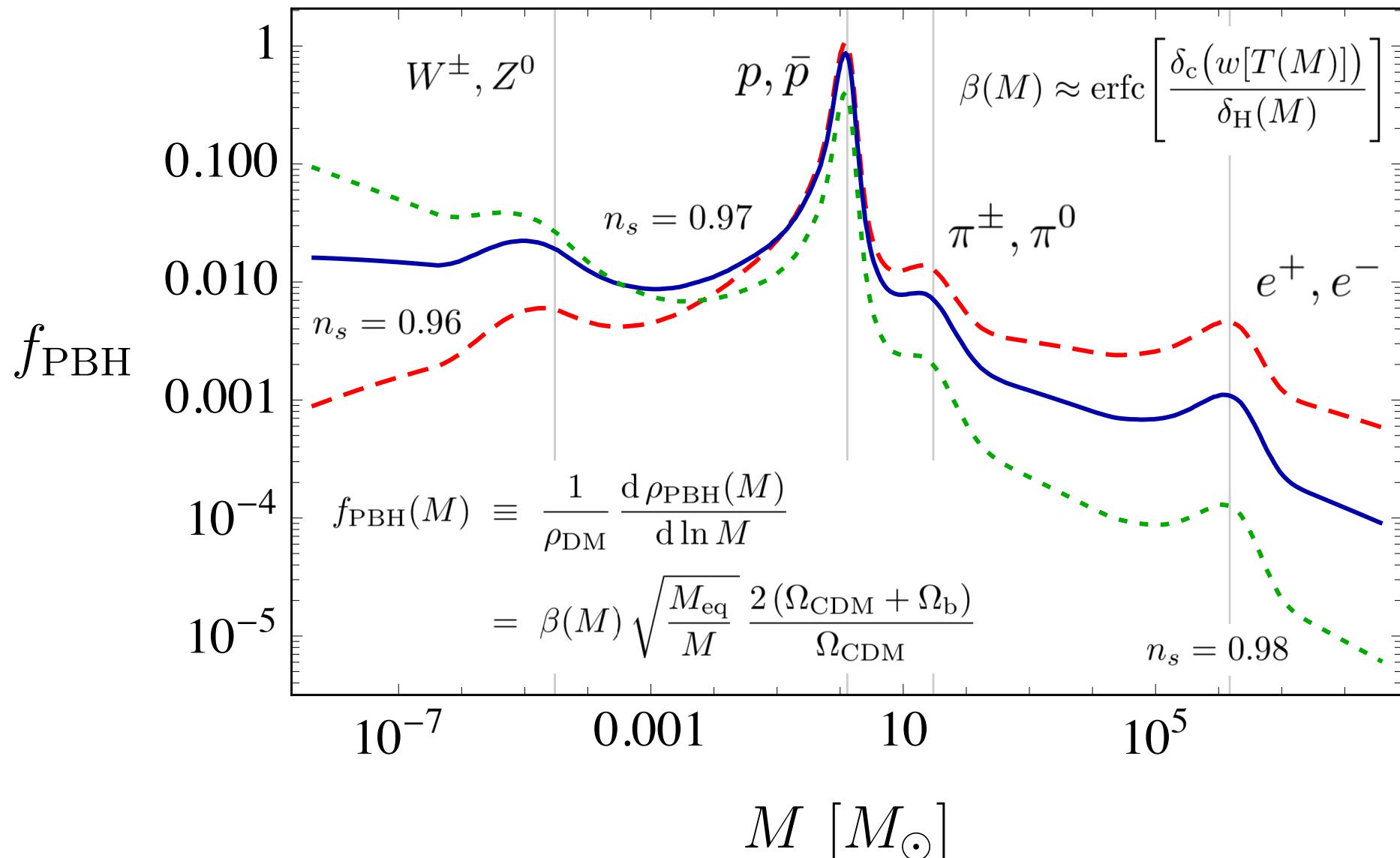


Thermal history of the universe

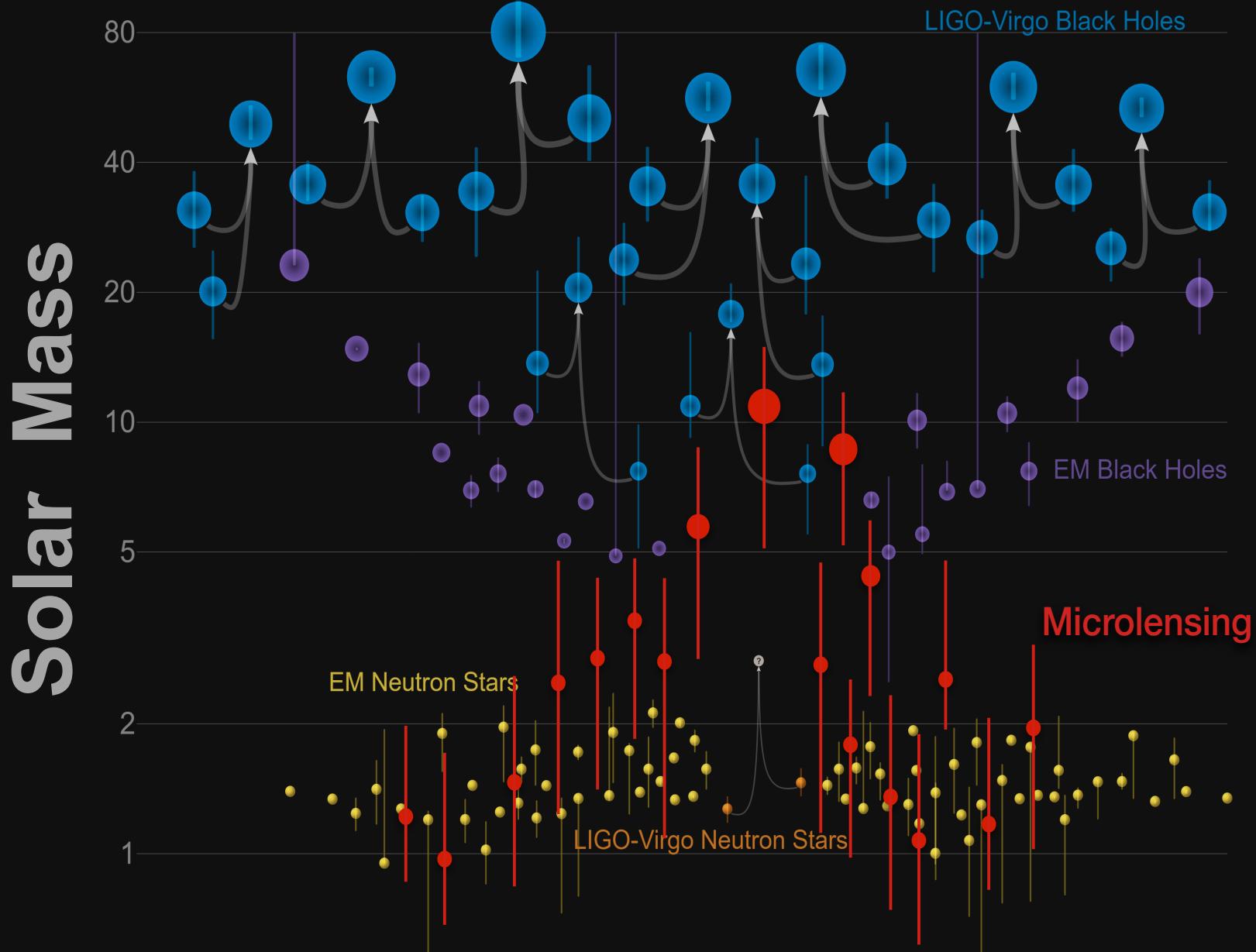


Predictions for PBH mass spectrum

Carr, Clesse, JGB, Kühnel (2019)

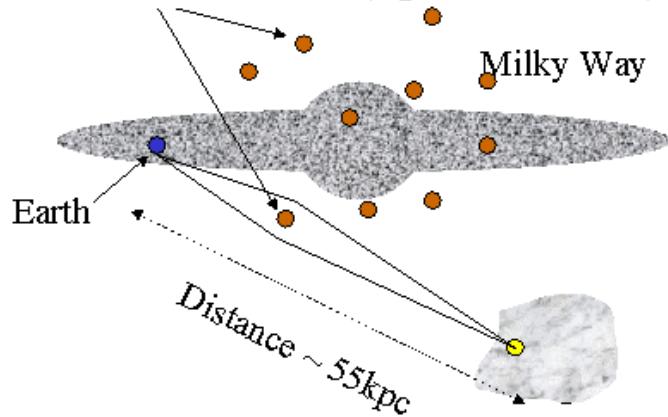


Black Holes and Neutron Stars

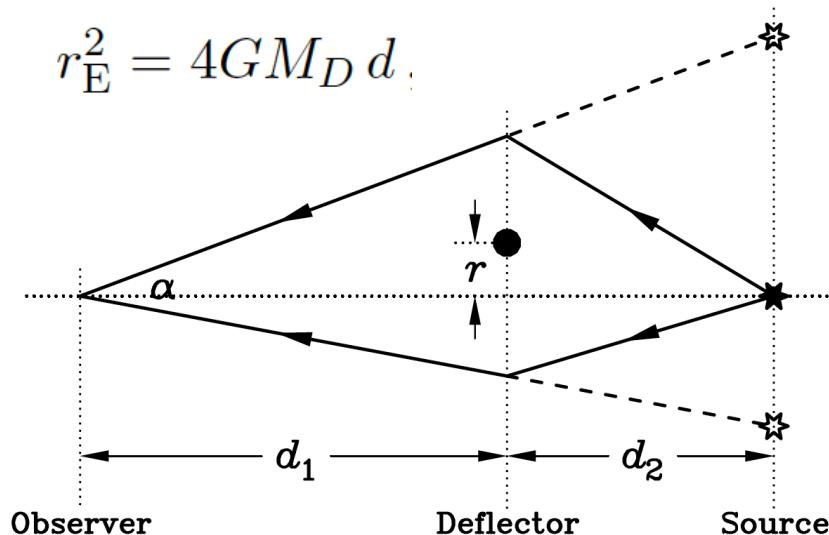


Microlensing Events

Gravitational lenses (e.g., brown dwarfs)



$$r_E^2 = 4GM_D d$$



$$d = \frac{d_1 d_2}{d_1 + d_2}$$

$$A = \frac{2+u^2}{u\sqrt{4+u^2}} \quad u = \frac{r}{r_E} \quad \text{amplification}$$

$$\overline{Dt} = \frac{r_E}{v} = \frac{\sqrt{4GM_D d}}{v} \quad \text{average } \frac{1}{2} \text{ crossing}$$

$$M_D = 100 M_{\odot} \Rightarrow \overline{Dt} = 4 \text{ years}$$

$$M_D = 10 M_{\odot} \Rightarrow \overline{Dt} = 1.23 \text{ years}$$

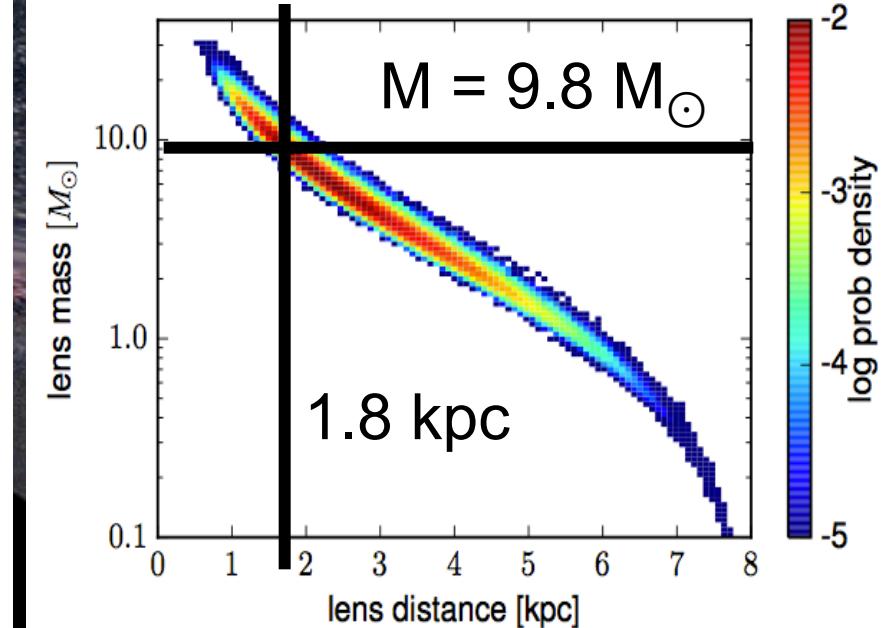
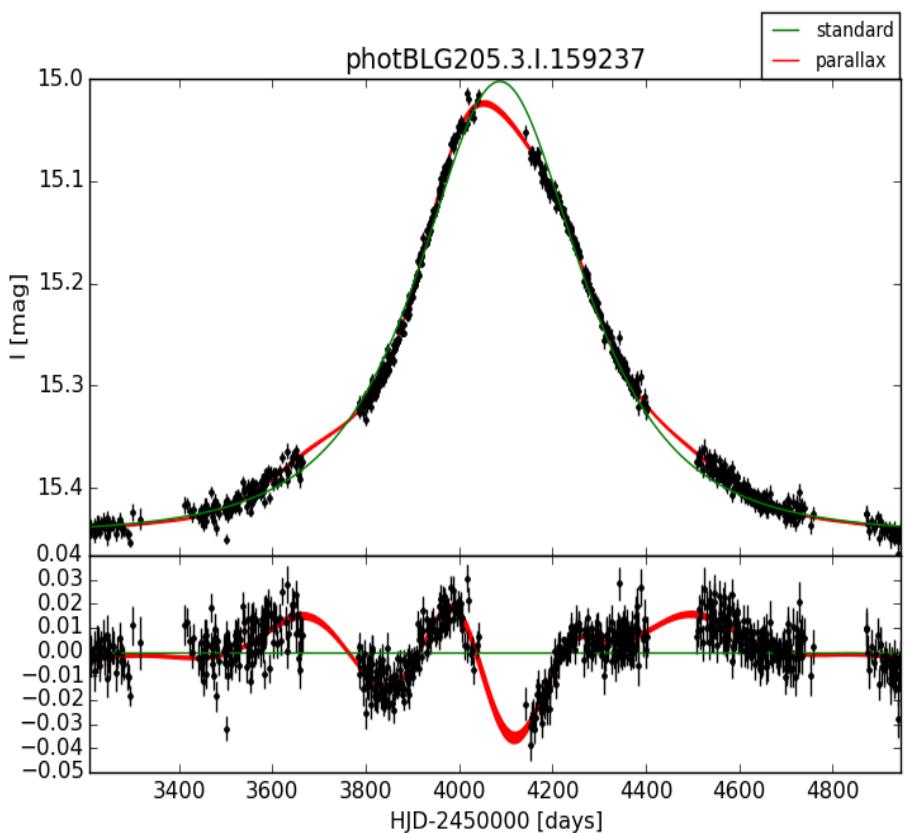
$$M_D = 1 M_{\odot} \Rightarrow \overline{Dt} = 5 \text{ months}$$

$$M_D = 0.1 M_{\odot} \Rightarrow \overline{Dt} = 1.5 \text{ months}$$

$$M_D = 0.01 M_{\odot} \Rightarrow \overline{Dt} = 2 \text{ weeks}$$

OGLE3-UL-PAR-02 - candidate BH

Wyrzykowski (2016)



OGLE photometry
from 2001-2008
and microlensing model

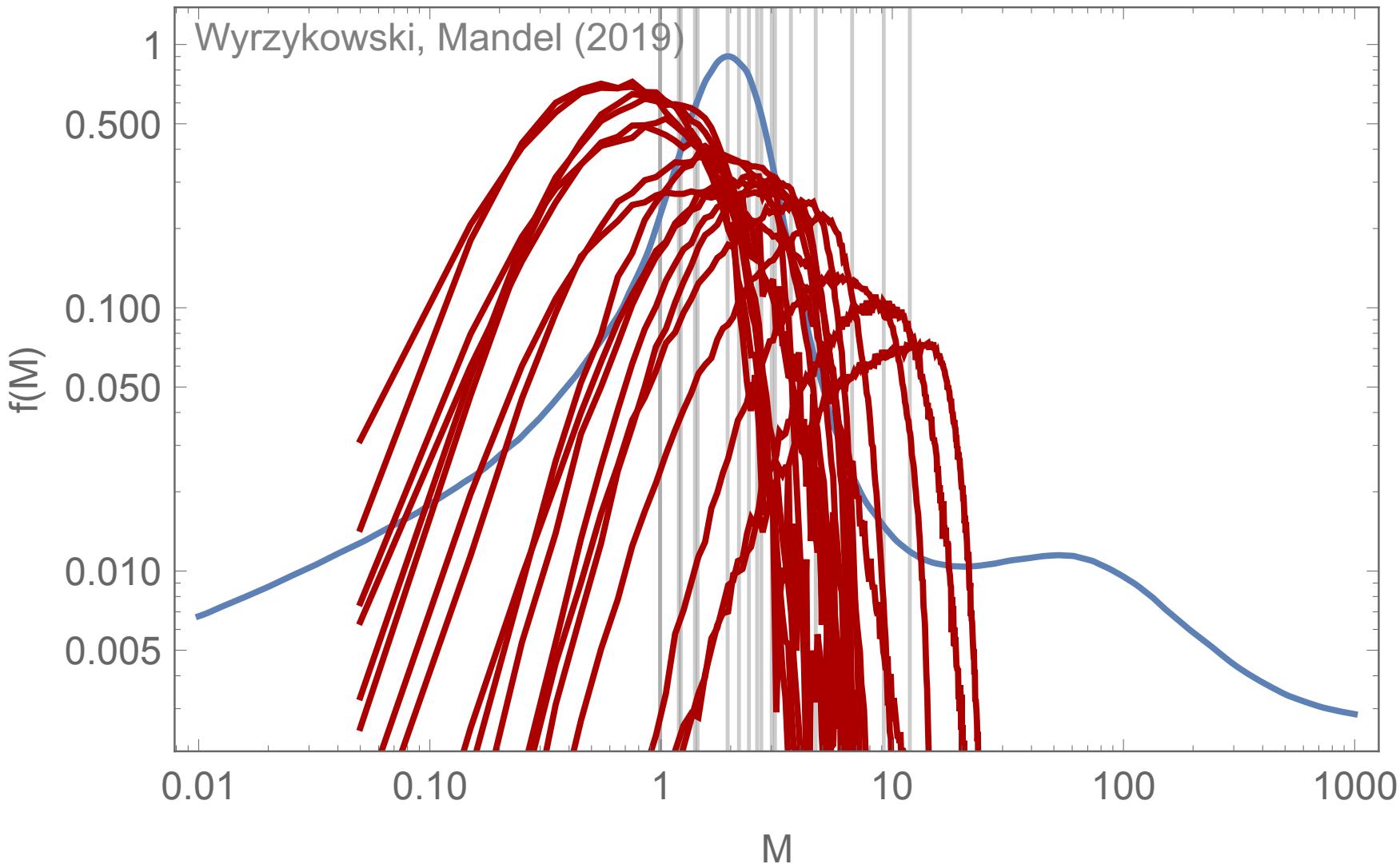


Mass, Distance

(degenerated estimate)

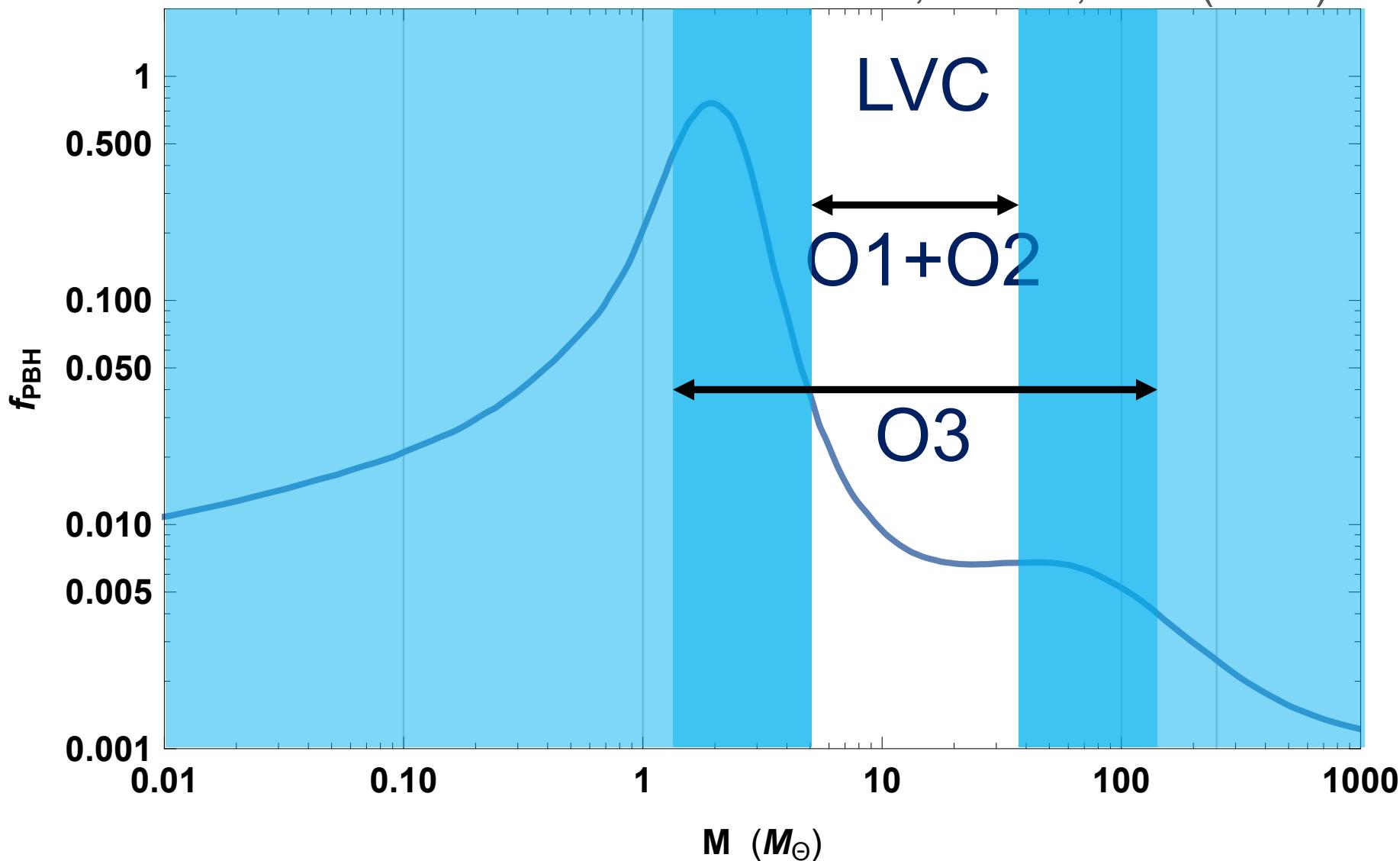
OGLE-GAIA microlensing events

Carr, Clesse, JGB, Kühnel (2019)

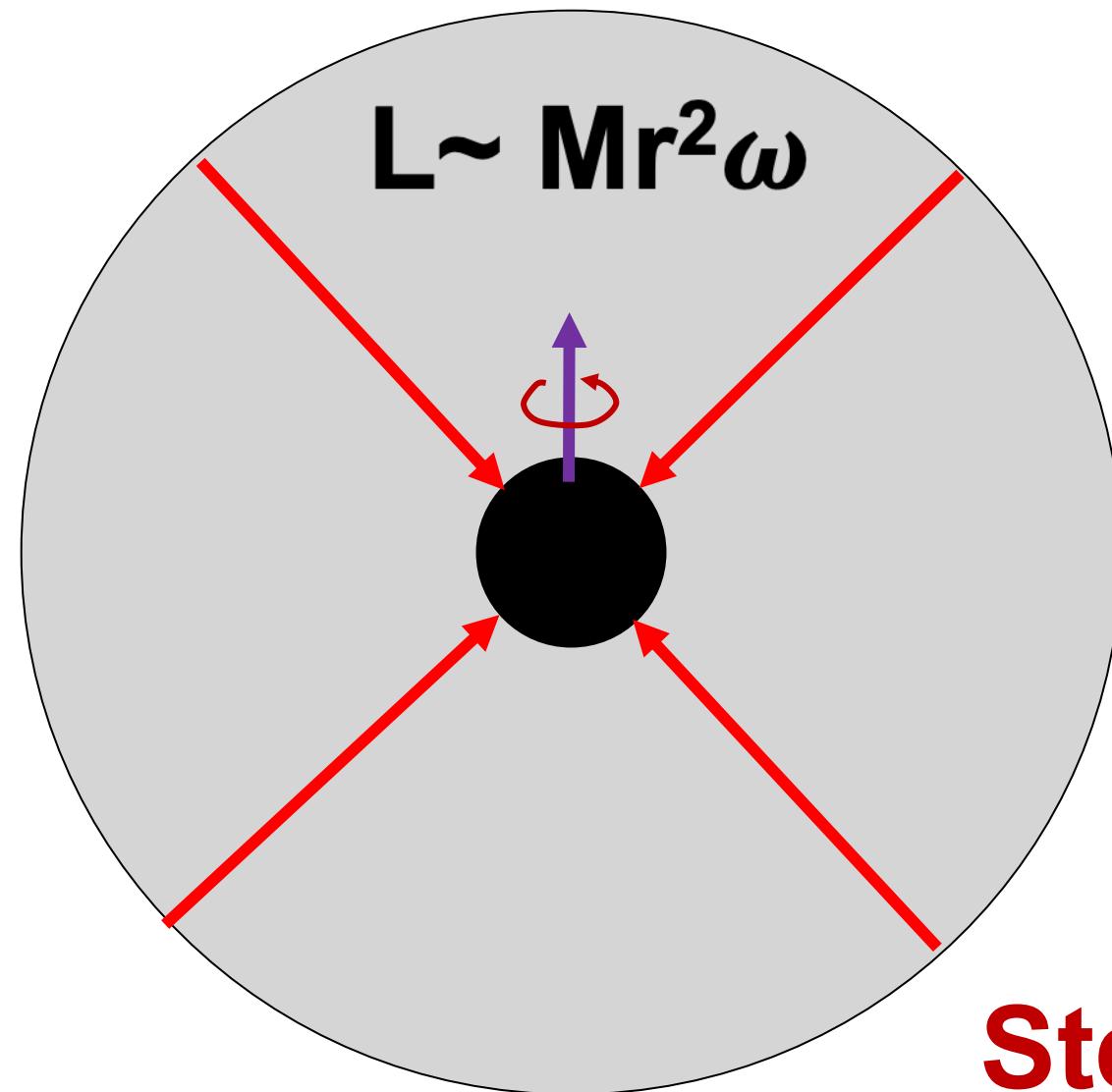


Predictions for PBH mass spectrum

Carr, Clesse, JGB (2019)

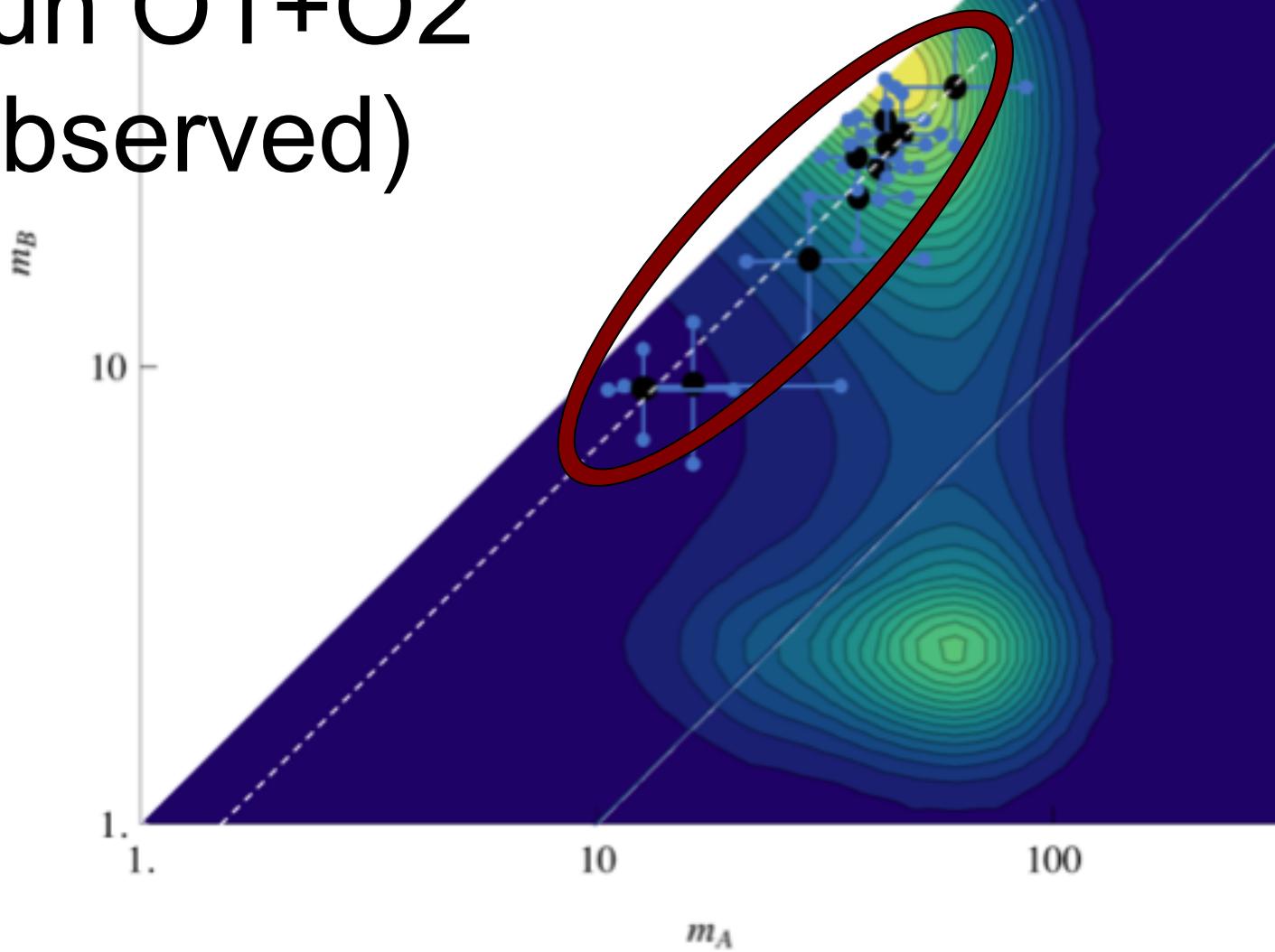


PBH are \sim spinless



Primordial
BH
= Mass
Stellar BH

Rate of merger events @ LIGO-Virgo Run O1+O2 (observed)



Carr, Clesse, JGB, Kühnel (2019)

Rate of merger events

@ LIGO-Virgo

Run O3

(predicted)

m_2

10

1.

1.

m_1

Carr, Clesse, JGB, Kühnel (2019)

